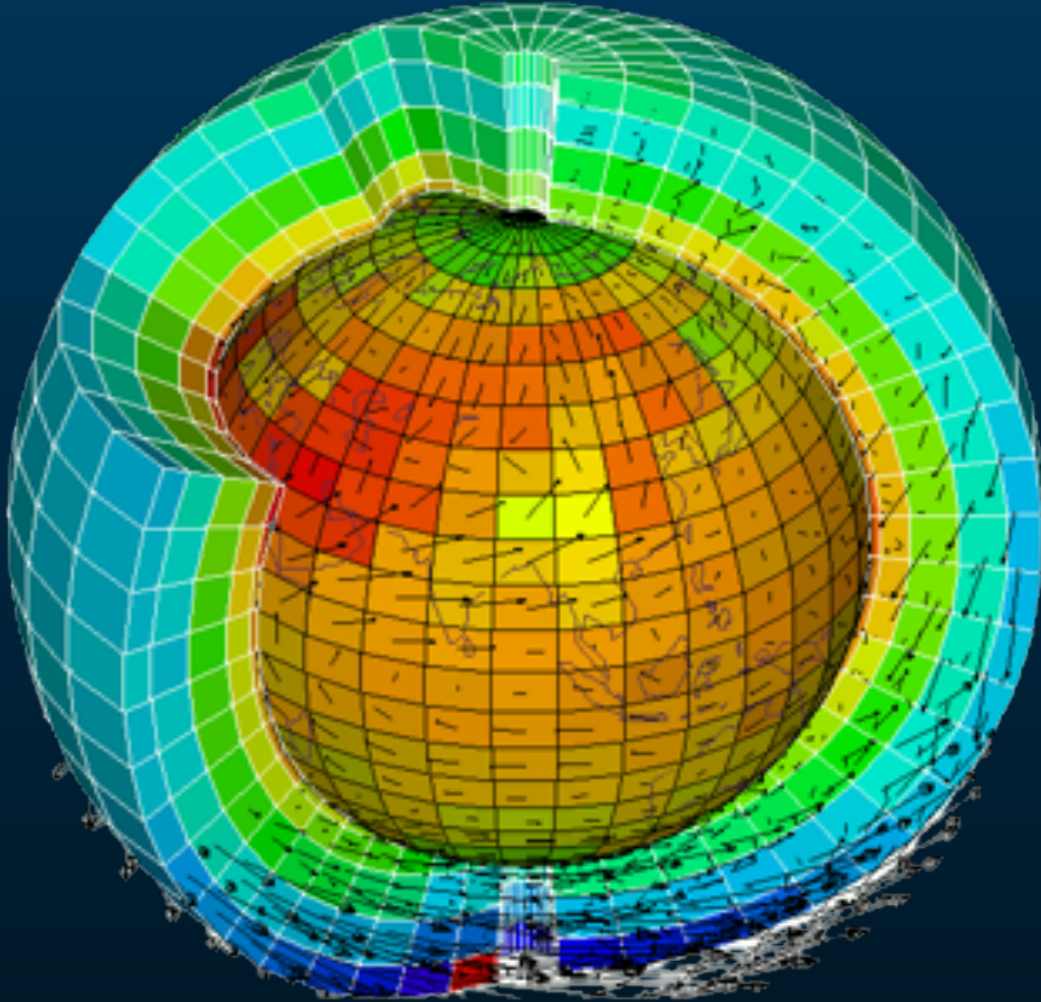


# Des données spectroscopiques à la modélisation d'exoplanètes

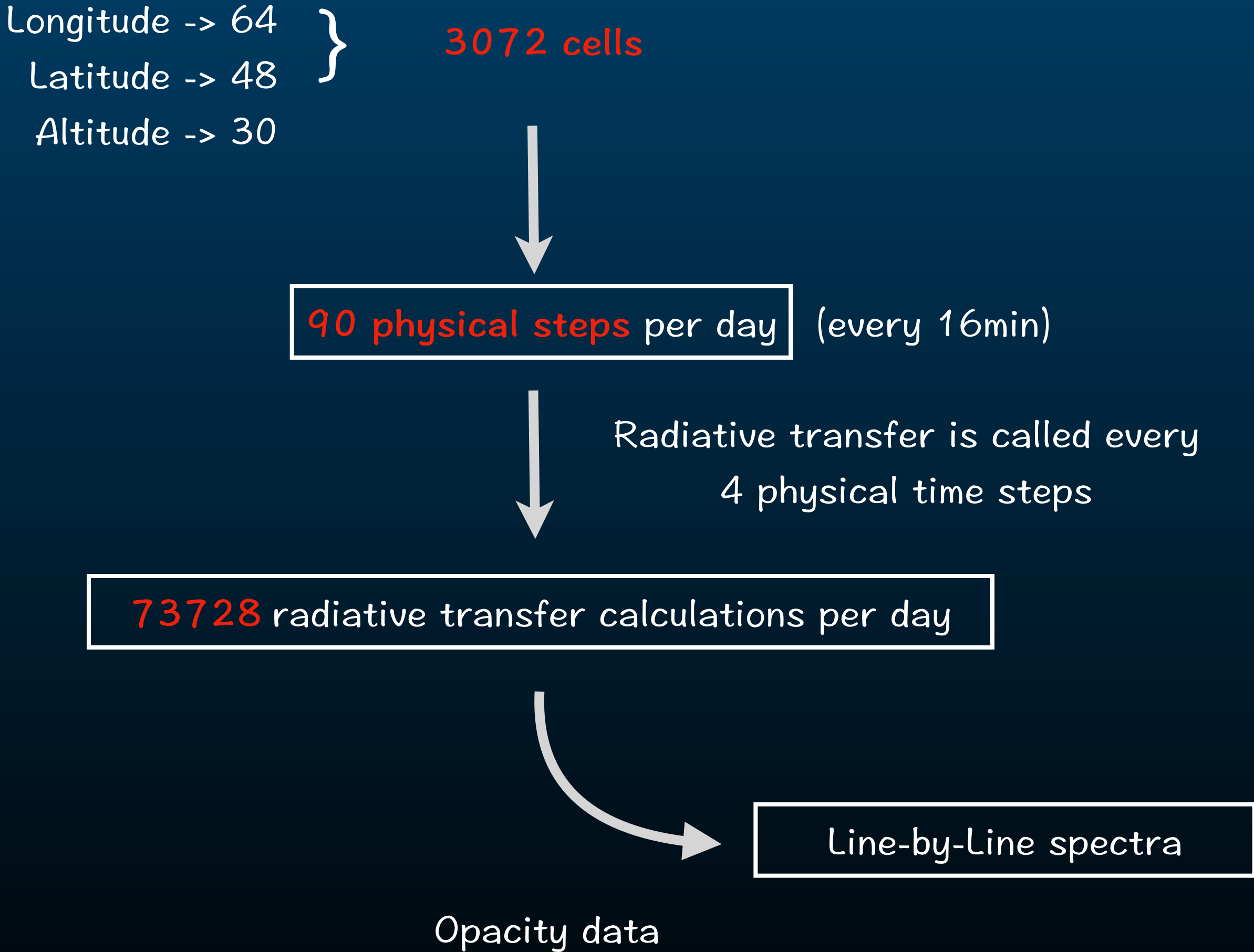
**Guillaume Chaverot**  
IPAG - University of Grenoble-Alpes

In collaboration with J.-M. Hartmann, M. Turbet, H. Tran, A. Campargue, D. Mondelain

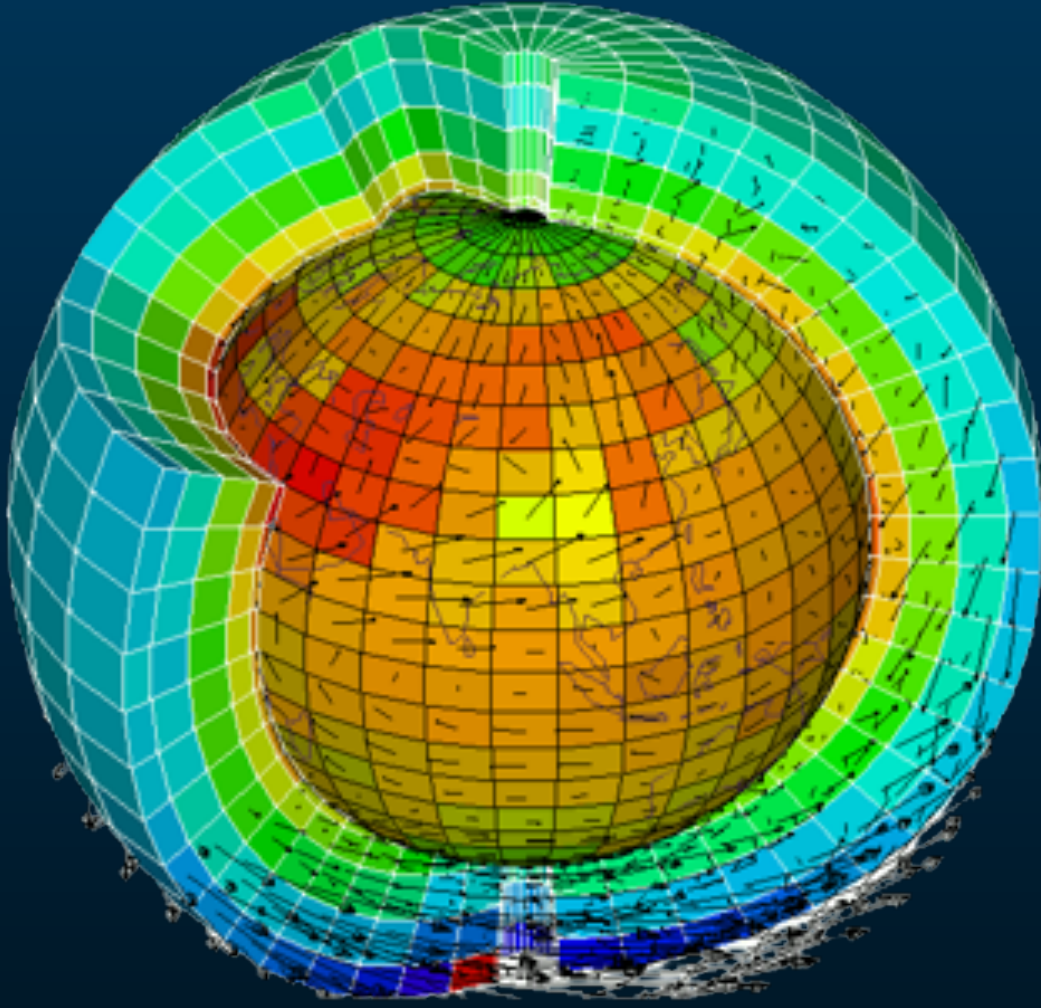
# Radiative transfer calculation



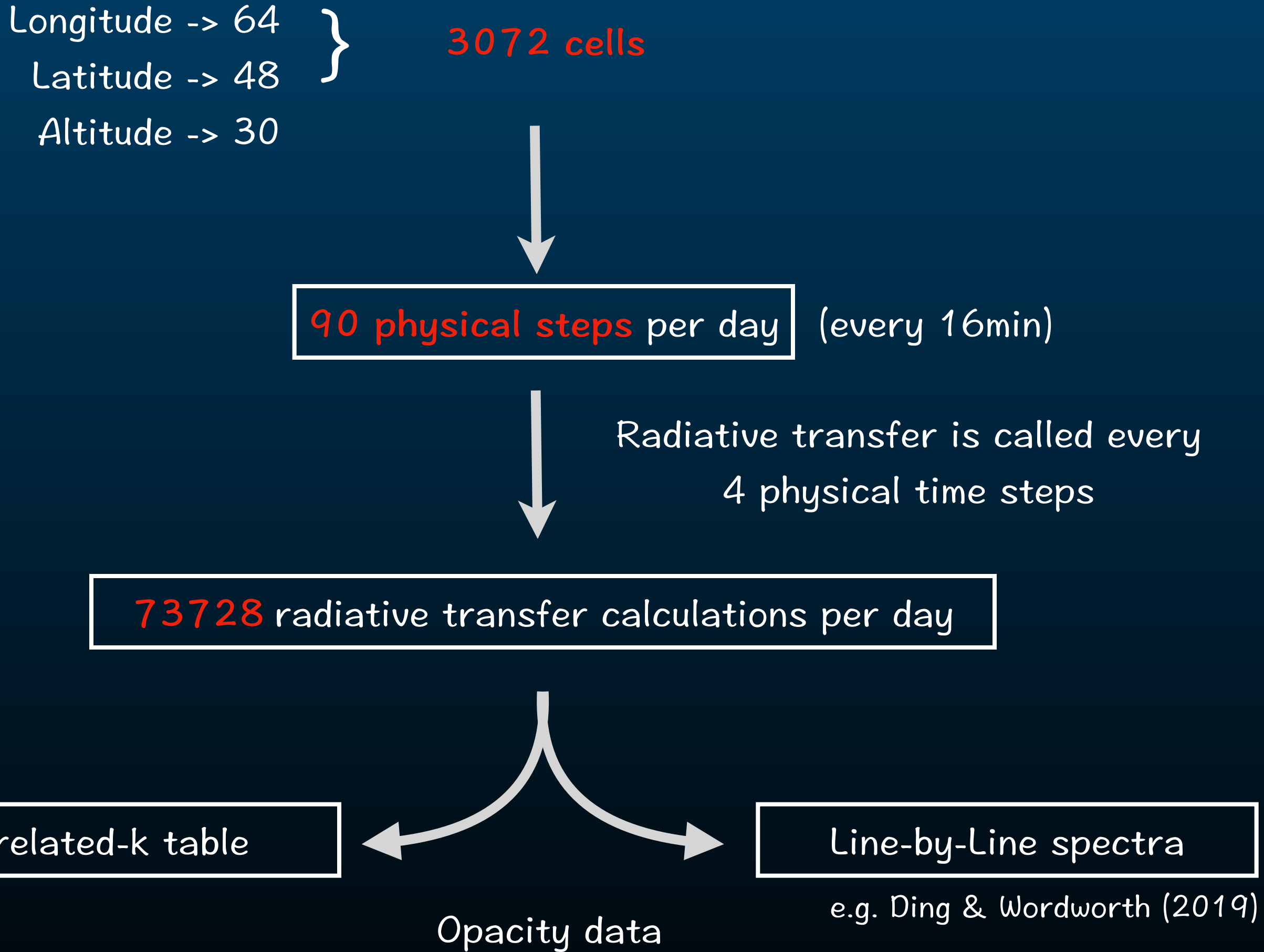
Generic-PCM



# Radiative transfer calculation



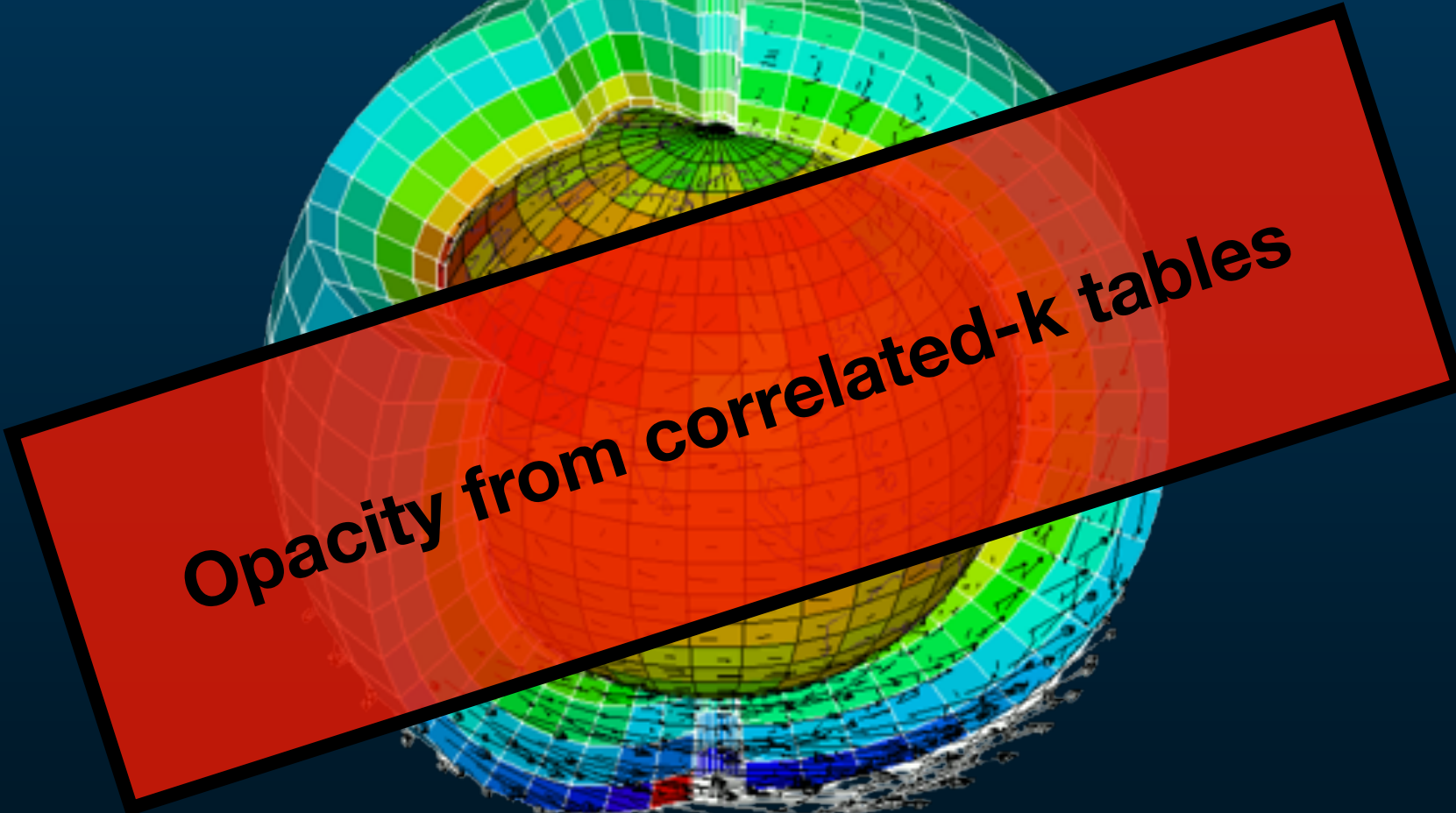
Generic-PCM



# Radiative transfer calculation



Longitude -> 64 }  
Latitude -> 48 } **3072 cells**  
Altitude -> 30 }



Generic-PCM

**90 physical steps** per day (every 16min)

Radiative transfer is called every 4 physical time steps

**73728** radiative transfer calculations per day

Correlated-k table

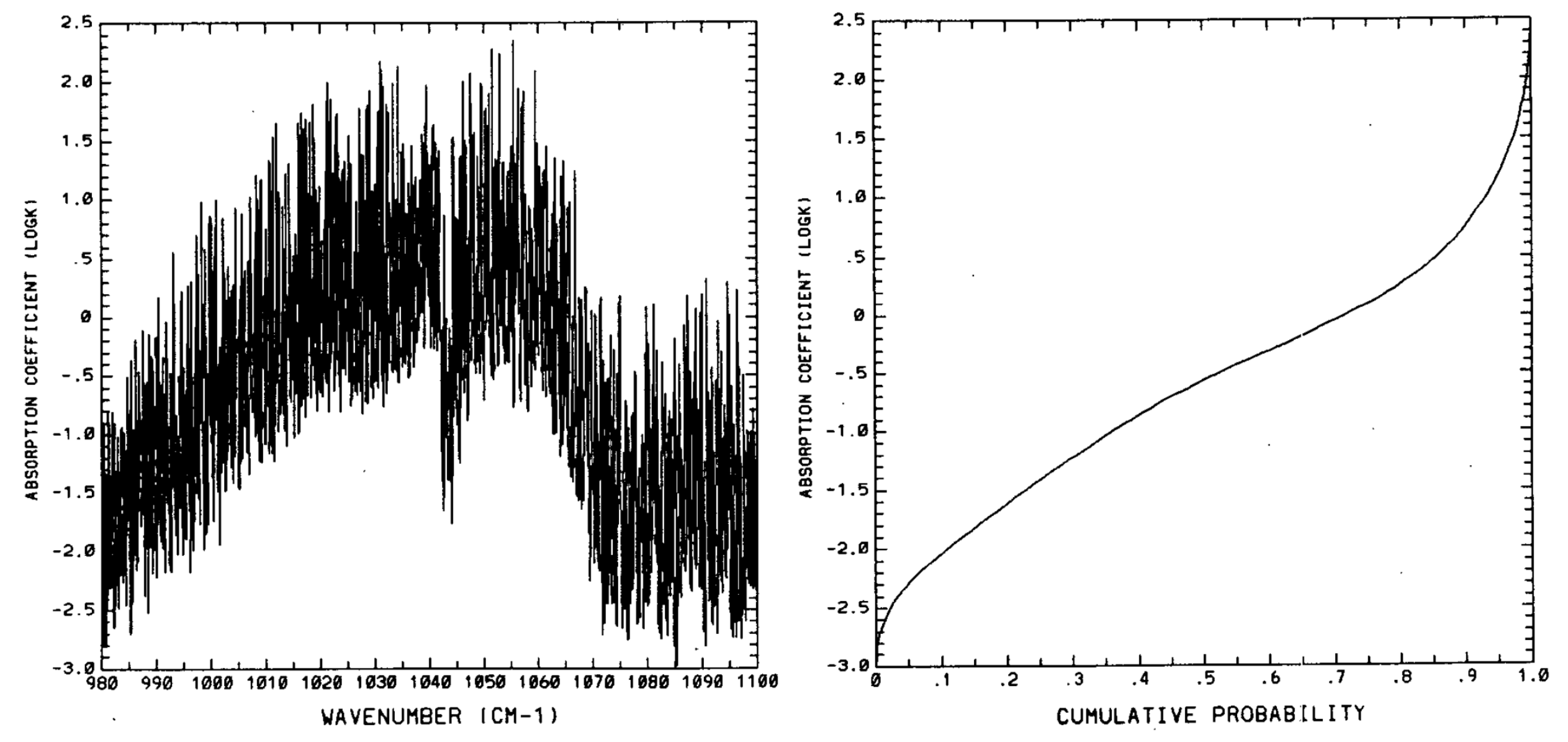
Opacity data

# Radiative transfer calculation



Opacity from correlated-k tables

Generic-PCM

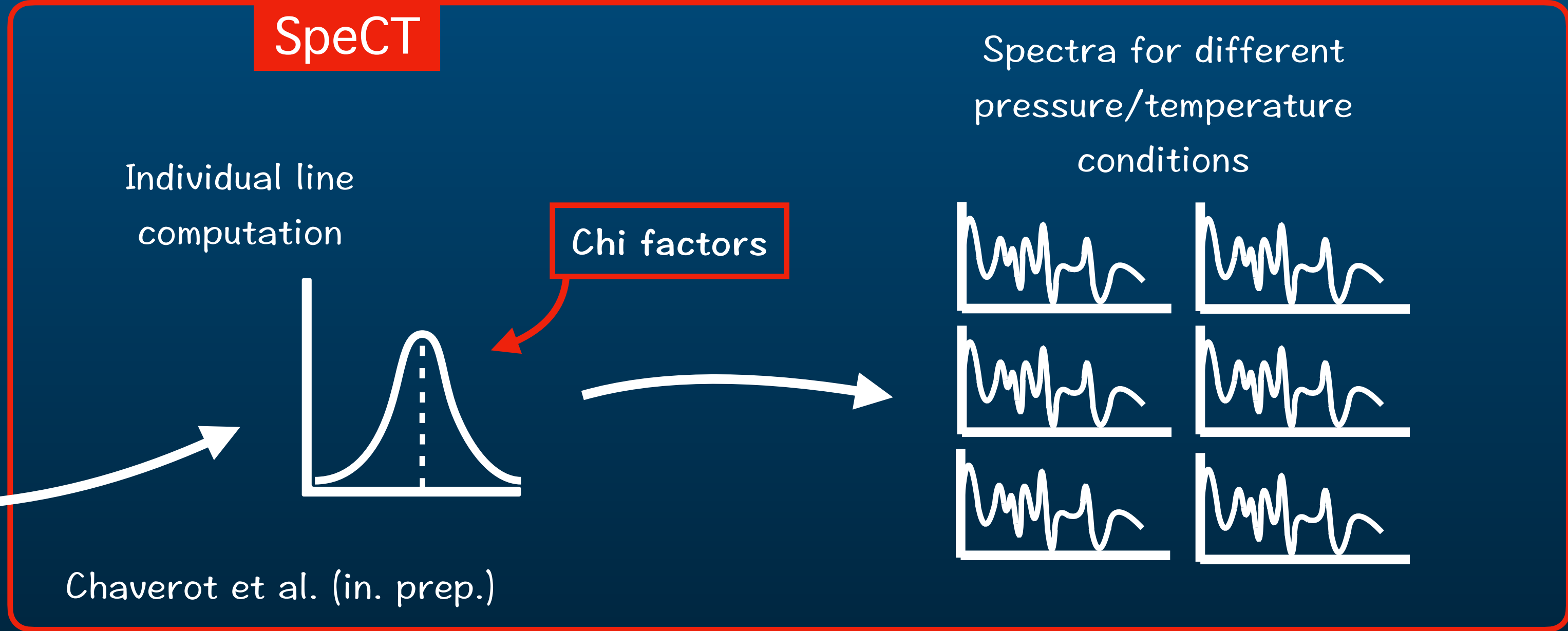


Fu et al. (1992)

Correlated-k table

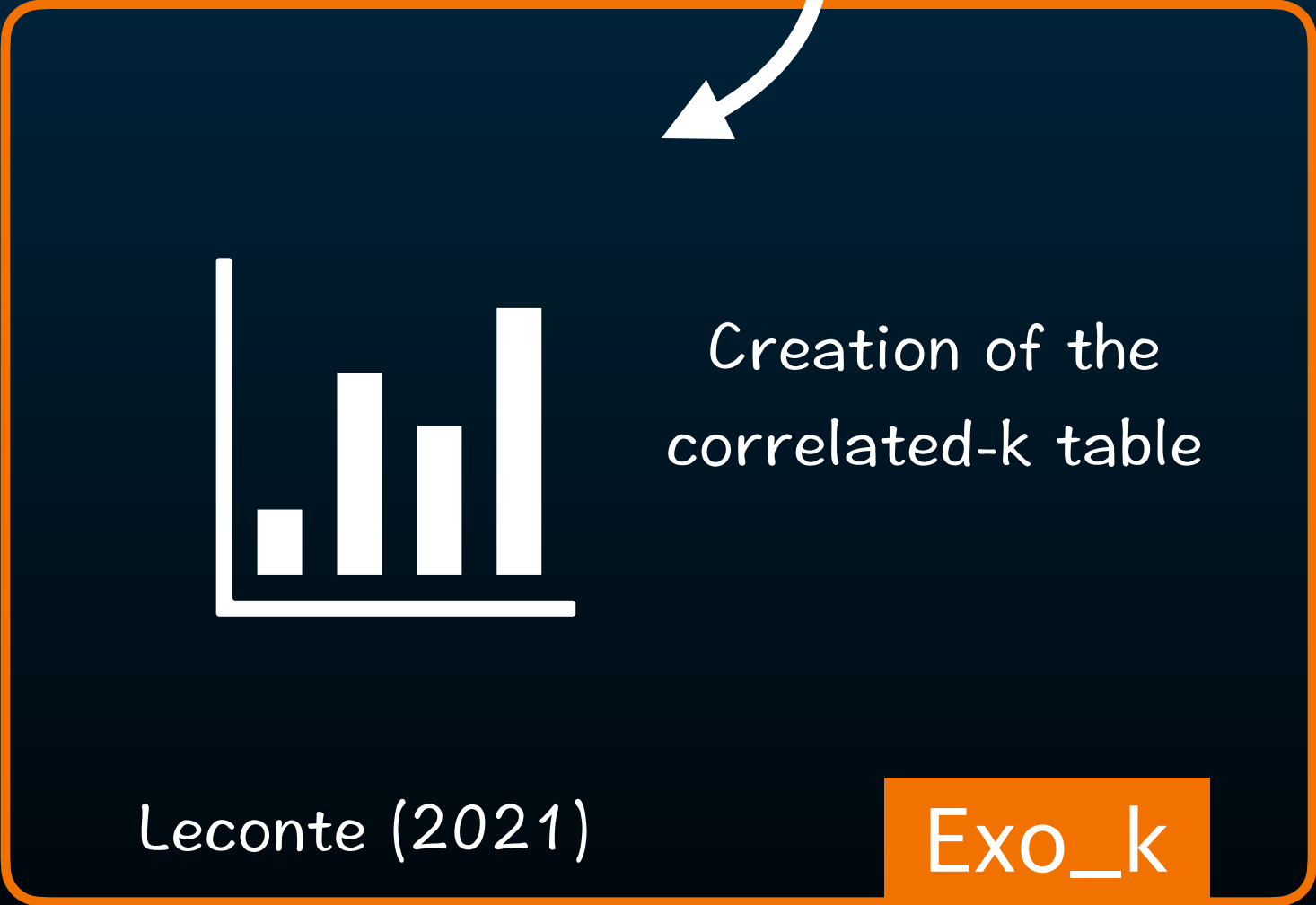
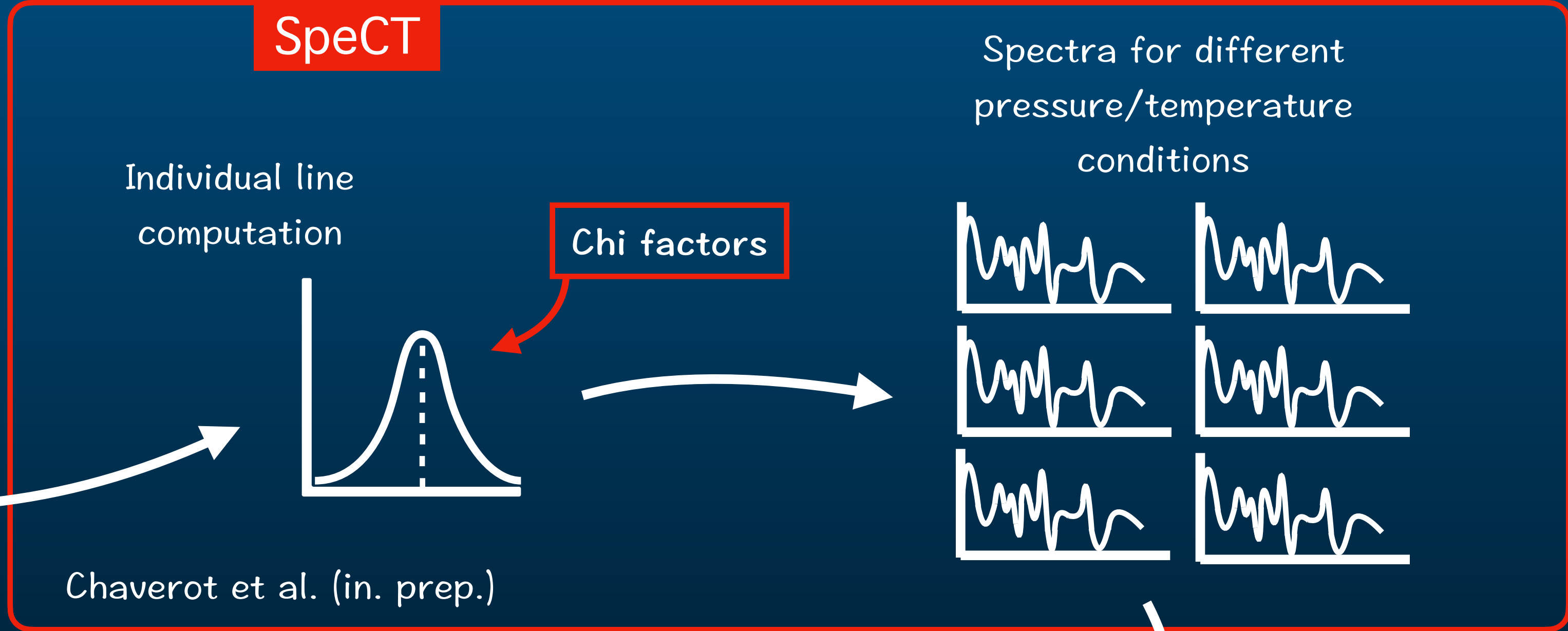
# Making opacity data

Spectroscopic line list  
(HITRAN, HITEMP ...)



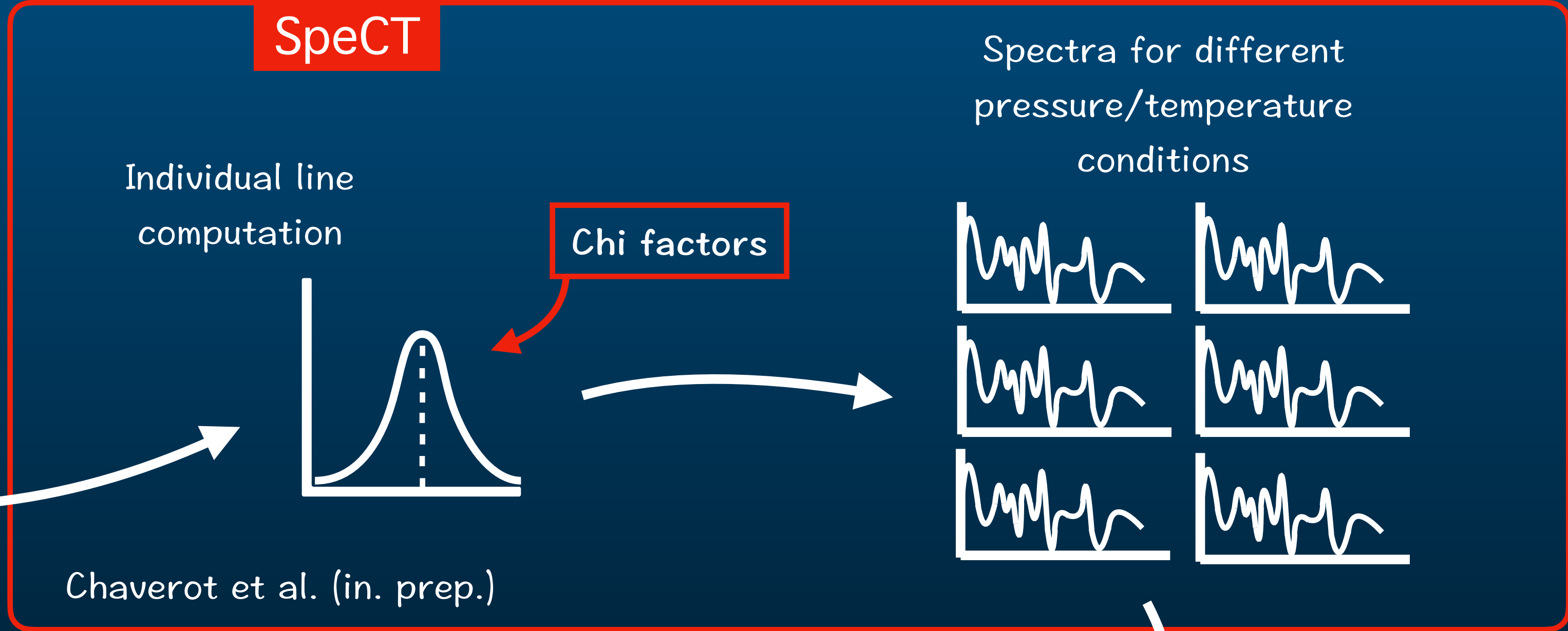
# Making opacity data

Spectroscopic line list  
(HITRAN, HITEMP ...)

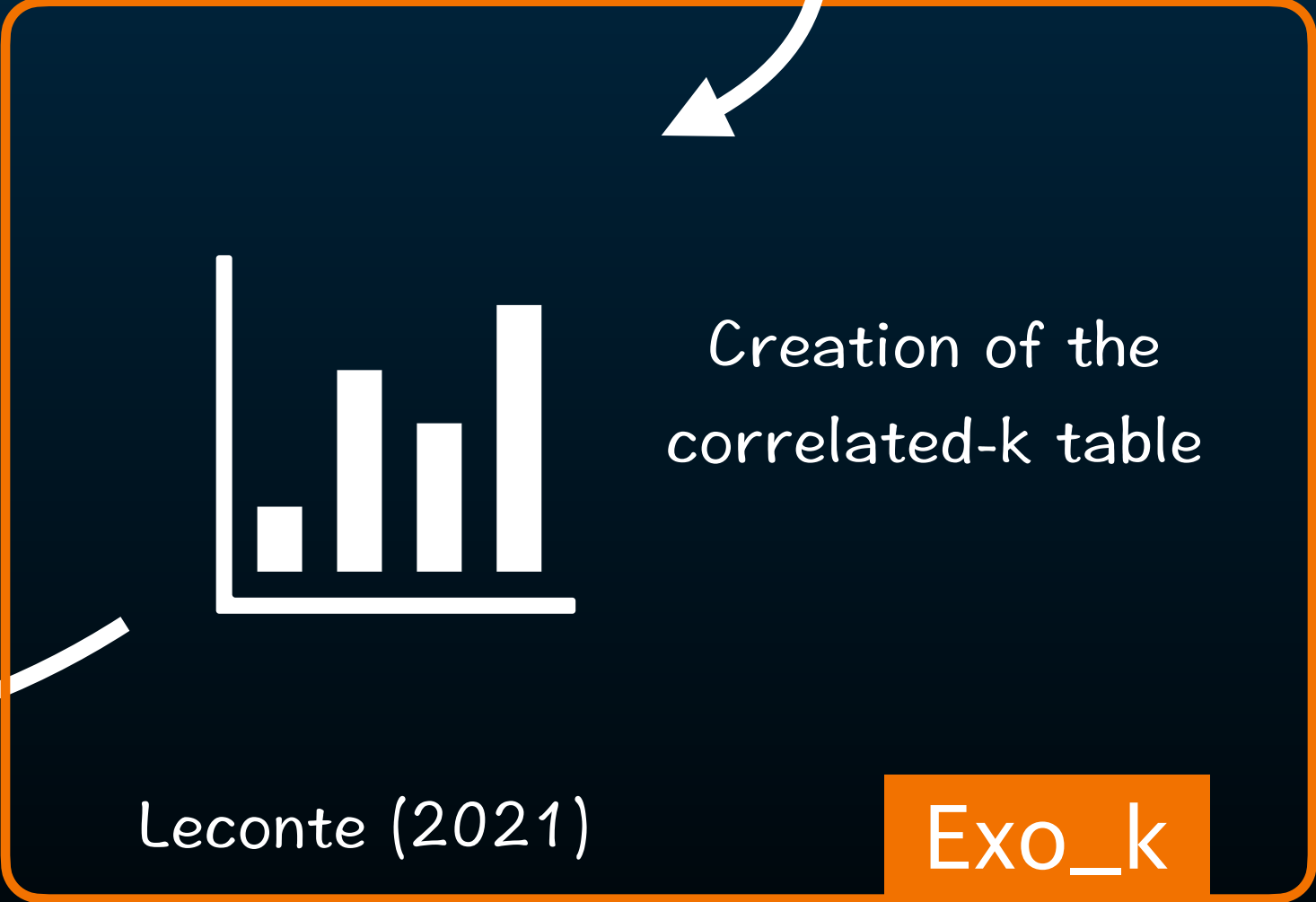
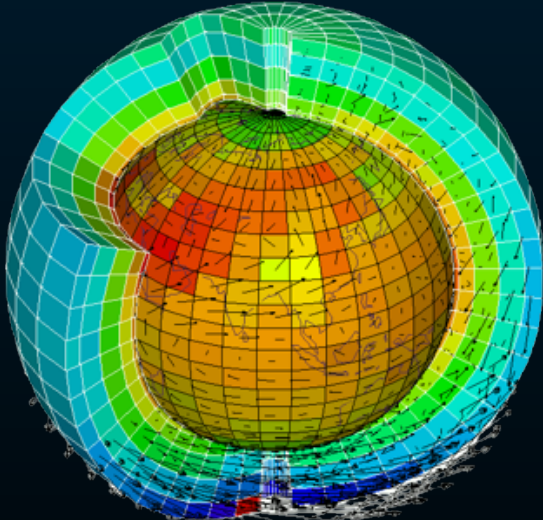


# Making opacity data

Spectroscopic line list  
(HITRAN, HITEMP ...)

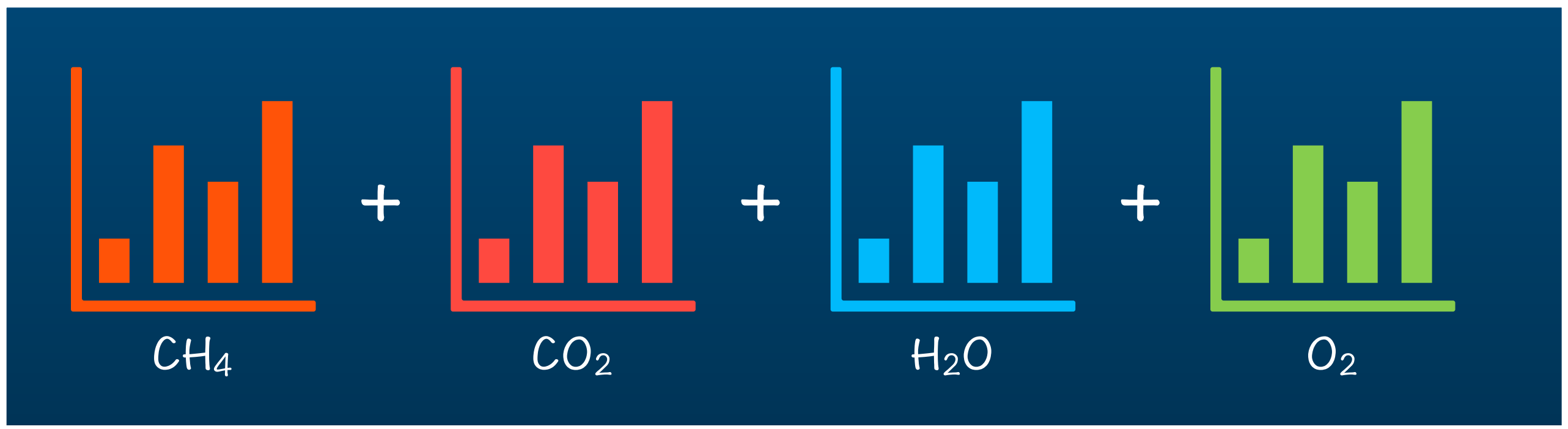


Climate modeling

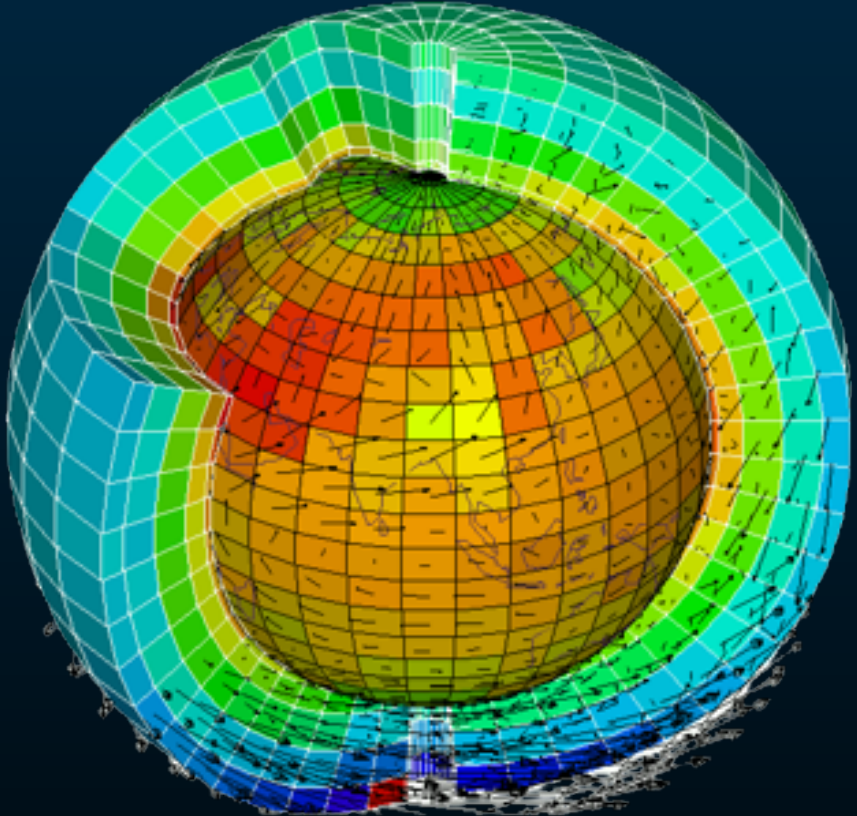




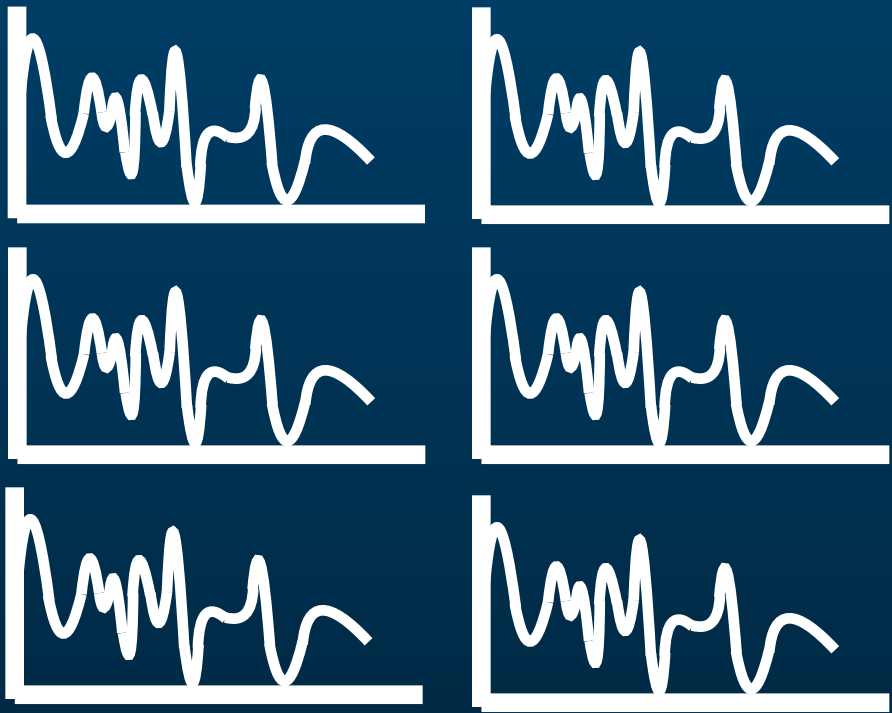
# Challenge of the composition



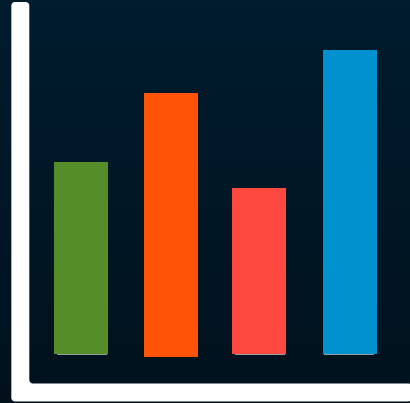
 No interpecies interaction



# Challenge of the composition



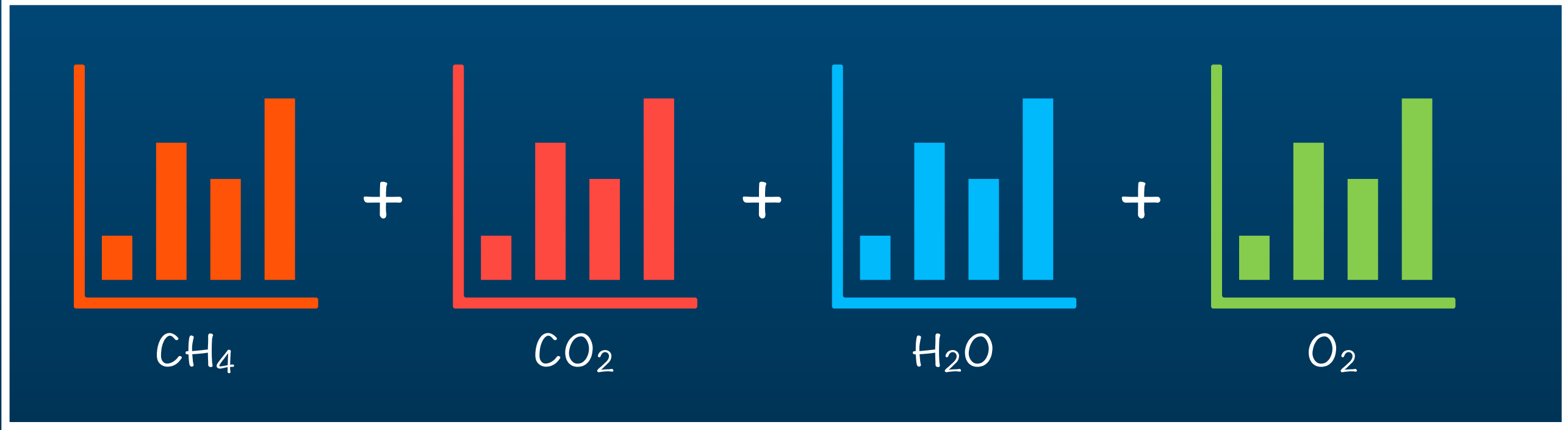
$H_2O + CO_2 + O_2 + CH_4$



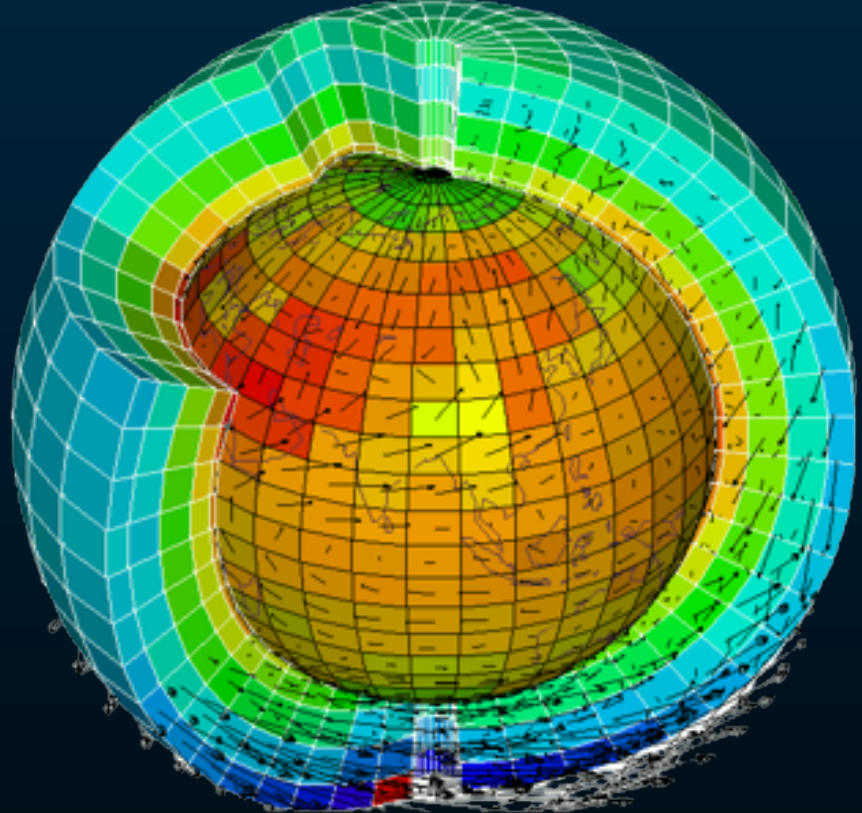
$H_2O + CO_2 + O_2 + CH_4$



Accurate opacities but one single composition -> time consuming



No interpecies interaction



# Spectra for Correlated-k Tables (SpeCT)

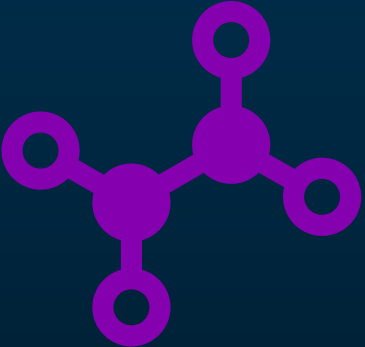


**PRESSURE**  
From few Pa to tens of bar

**TEMPERATURE**  
From few Kelvin to few thousands  
  
0K > HITRAN > 400K  
400K < HITEMP < 3000K



**Various gas mixtures**  
  
H<sub>2</sub>O+N<sub>2</sub>  
H<sub>2</sub>O+CO<sub>2</sub>  
CO<sub>2</sub>+N<sub>2</sub>  
H<sub>2</sub>O+CO<sub>2</sub>+N<sub>2</sub>

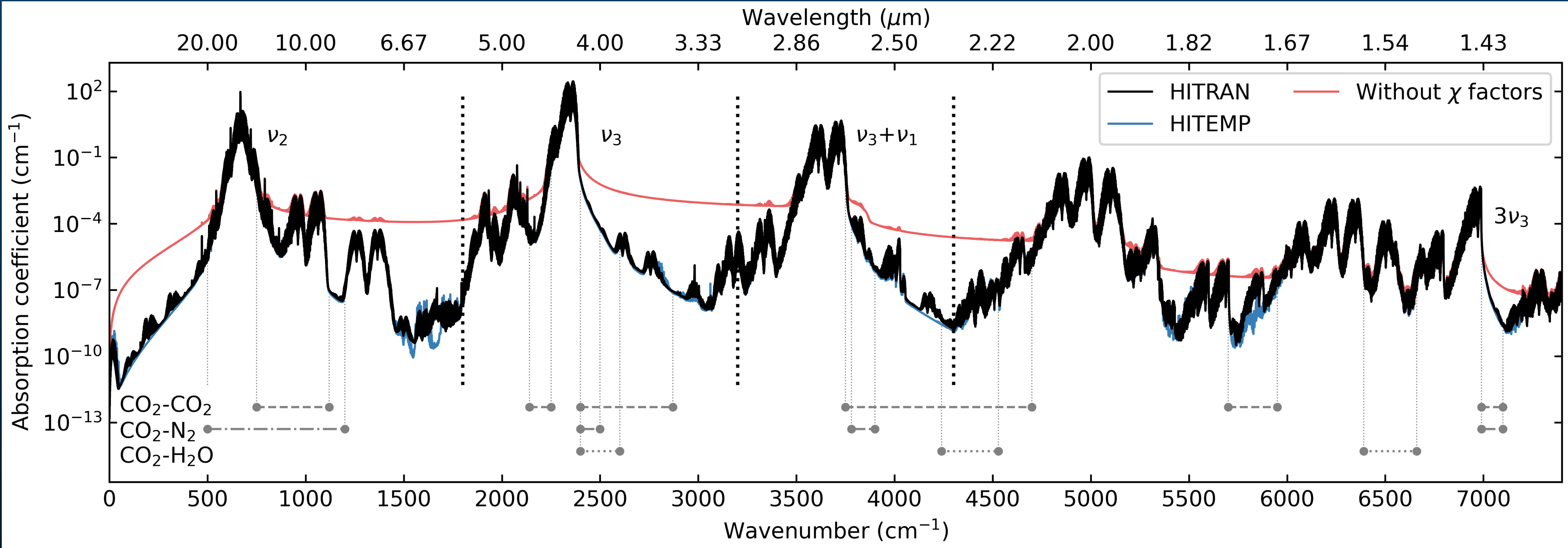


**WAVELENGTH**  
Visible and infrared up to microwave  
350nm -> 10mm  
  
High resolution to resolve each individual line



# Correction factors

Chaverot et al. (in. prep.)



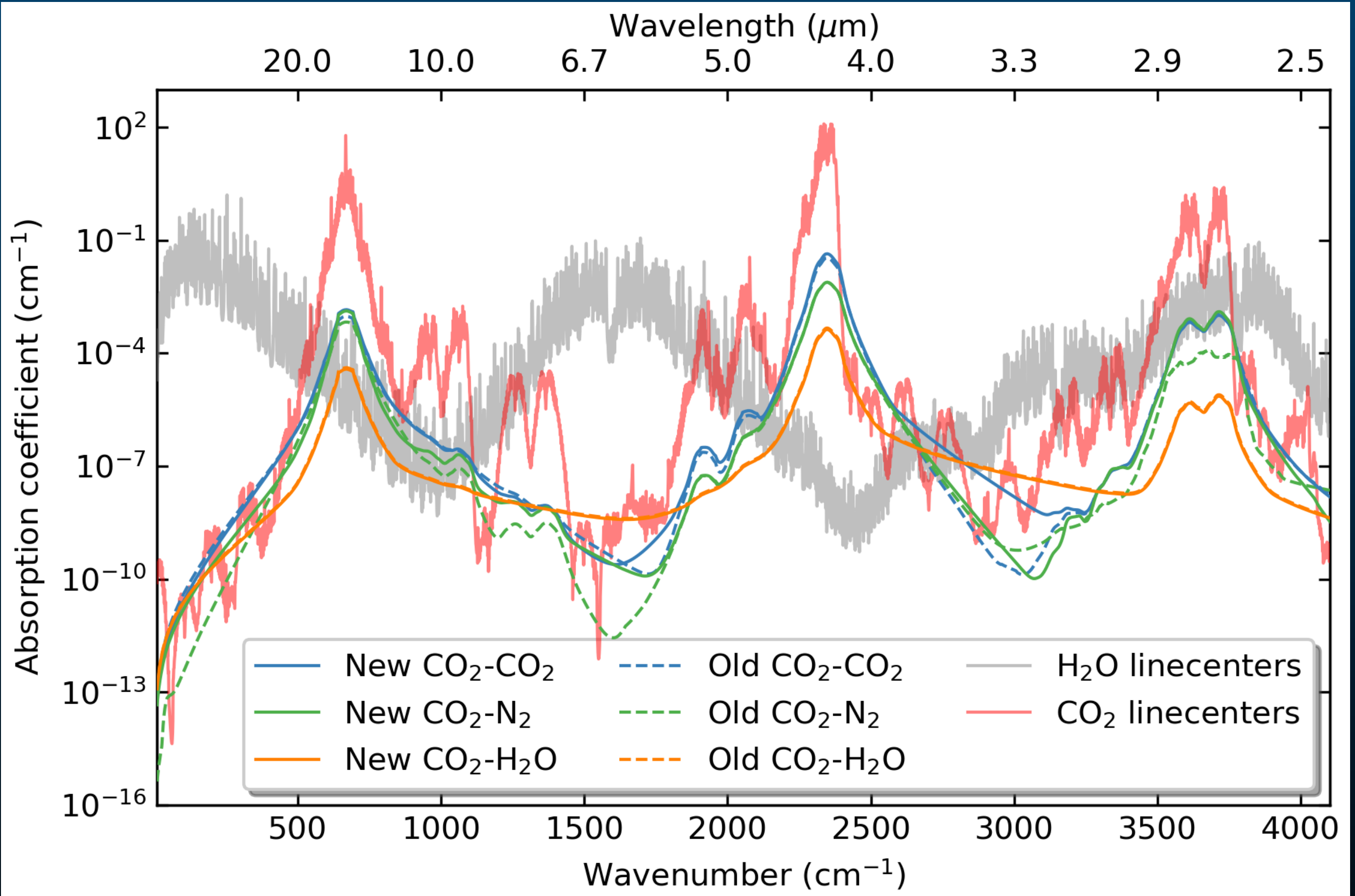
**Aims:**

- 1) Propose accurate opacity data for the climate modeling community
- 2) Calculate CO<sub>2</sub> continua, based on the Chi factor formalism

# Correction factors

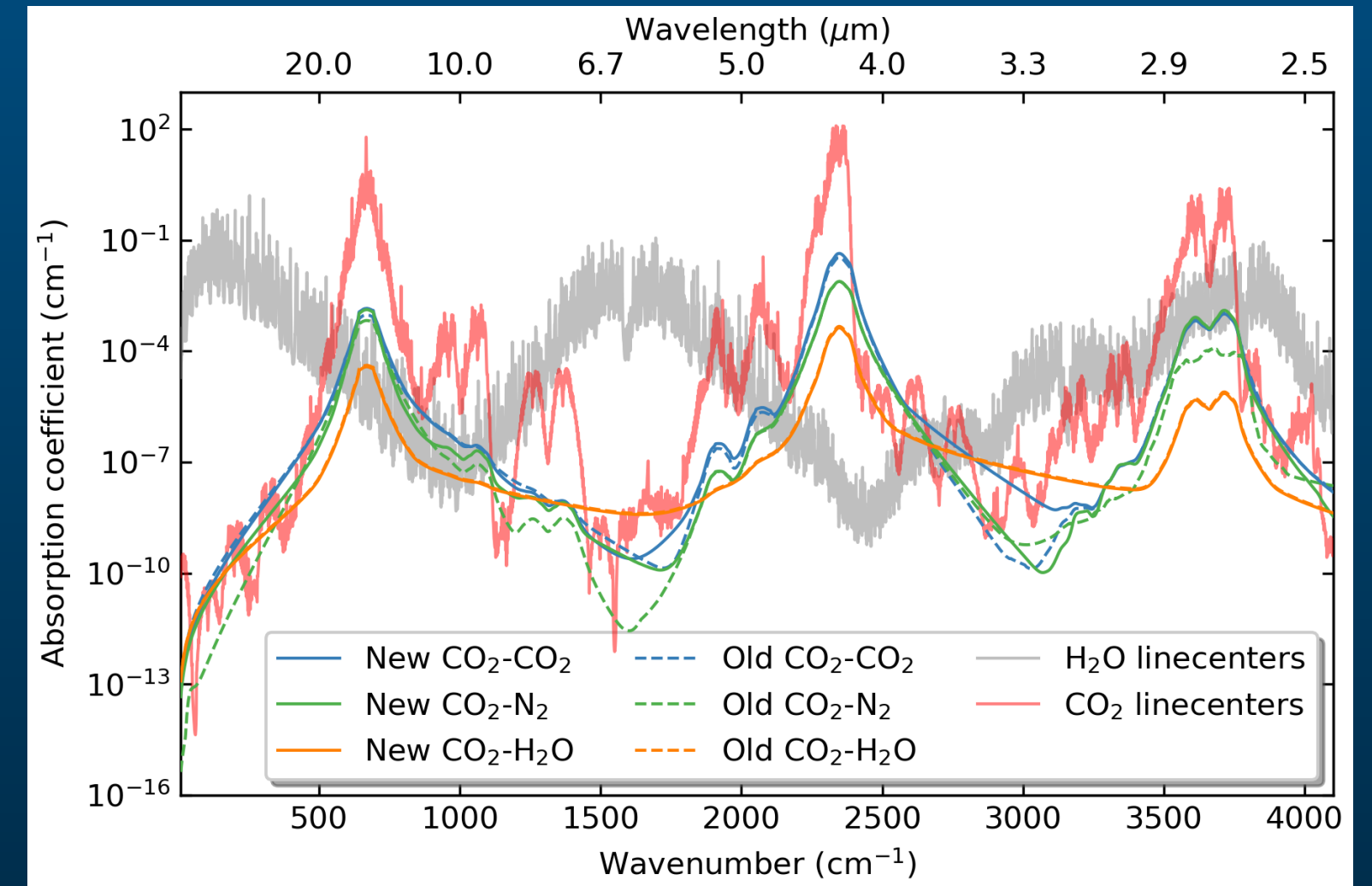
Updated Chi factors:

- common formalism (Perrin & Hartmann 1989)
- using HITRAN 2020
- new experiment data: pure CO<sub>2</sub>, 2600-2900cm<sup>-1</sup> (Tran et al. 2024)
- Chi factor for a new band:  $\nu_2$ , CO<sub>2</sub>-N<sub>2</sub> (Niro et al. 2004)



Chaverot et al. (in. prep.)

# Next steps



## Development of SpeCT:

- adding other greenhouse gases (e.g. CH<sub>4</sub>)
- adding inter-species broadening for more gases (e.g. H<sub>2</sub>, O<sub>2</sub>)
- adding traces gas (CO, SO<sub>2</sub>, He, H<sub>2</sub>S)

## Long term aims:

- creating an open-source database of correlated-k tables using SpeCT (project coordinated by M. Turbet)
- producing continua for all the molecules that are relevant for temperate exoplanets